

Optical diagnostics in fluid mechanics

Metrology of particles

Part 1 : Fundamentals

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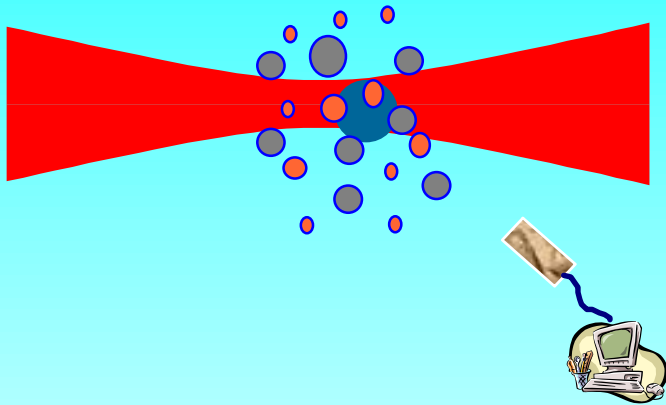
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Outline of course



- Elastic scattering
- Non-elastic scattering

- Simple scattering
- Multiple scattering
- Cohérente scattering

- Introduction
- Fundamentals
- Macroscopic models of light scattering by particles
 - **Approximate models**
 - **Rigorous theories**
 - **Numerical methods**
- Measurement techniques
 - LDV, PDA
 - Imaging techniques
 - Refractometry
 - Extinction spectrometry
- Research topics

Introduction – macroscopic models

➤ Approximate models

- Rayleigh: $l \ll \lambda$
- Diffraction: $l \sim \lambda$
- Geometrical optics: $l \gg \lambda$
- **VCRM – Vectorial Complex Ray Model**
- Rayleigh-Gans: $|m - 1| \ll 1$
- GTD - Geometrical Theory of Diffraction,

➤ Rigorous theories

- **TLM – Lorenz-Mie Theory**
- GLMT – Generalized Lorenz-Mie Theory
- Debye theory / series

➤ Numerical methods

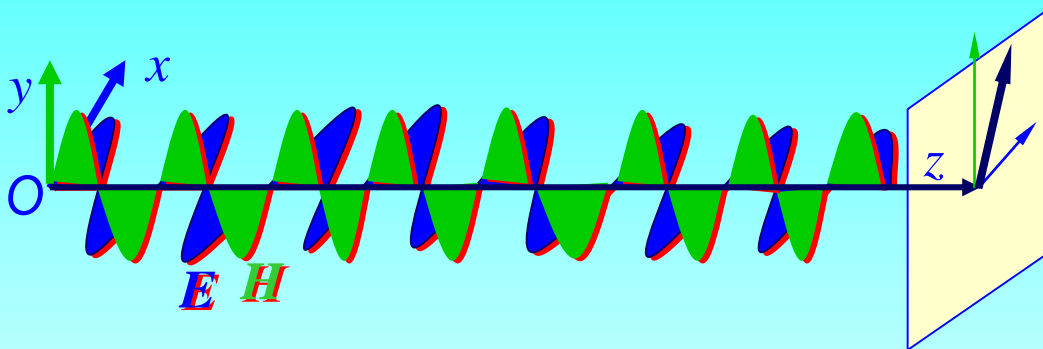
- T-Matrices,
- DDA - Dipole discretization Approximation (ADDA, DDSCAT)
- FDTD - Finite-difference time-domain
- MoM – Method of Moment
- FEM – Finite Element Method

Introduction – Measurement techniques

- **ADL – PDA:**
 - ✓ *velocity, size*
 - ✓ *individual particle*
- **Imaging techniques (holography, PIV, HPIV, ...):**
 - ✓ *Size, velocity*
 - ✓ *individual particle or cloud*
- **Rainbow refractometry:**
 - ✓ *Size, refractive index (very accurate)*
 - ✓ *Big particles*
- **Extinction spectrometry**
 - ✓ *Size, concentration*
 - ✓ *Small particles ($d \sim \lambda$)*

Fundamentals – plane wave

Electromagnetic wave (EM) and its properties



Wave front = plans \parallel xOy
Wave vector $k \perp E$ et H

Wave vector: k

Wave number: $k = \frac{2\pi}{\lambda}$

Two polarisations:
$$\begin{cases} E_x = A_x \cos(\omega t - \mathbf{k} \cdot \mathbf{r} - \phi_0) \\ E_y = A_y \cos(\omega t - \mathbf{k} \cdot \mathbf{r} - \phi_0) \end{cases}$$

$E = E_0 e^{i(\omega t - \vec{k} \cdot \vec{r} - \phi_0)}$ direction of $E \rightarrow$ polarization

In a homogeneous and isotropic medium:

$$\mathbf{D} = \varepsilon \mathbf{E}, \quad \mathbf{B} = \mu \mathbf{H} \quad \mathbf{H} = \frac{1}{\mu \omega} \mathbf{k} \wedge \mathbf{E}$$

E – electric field

H – magnetic field

ε – permittivity

μ – permeability

Fundamentals – complex refractive index

Definition and physical interpretation of refractive index

Complex refractive index : $\tilde{m}^2 = \varepsilon$

$$\tilde{m} = m_r - m_i \cdot i$$

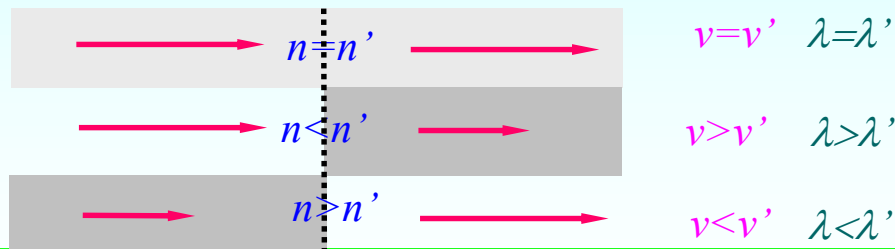
Real part – velocity : $m_r = \frac{c}{v}$

Examples:

In the vacuum (air): $c = 3 \times 10^8 \text{ m} \cdot \text{s}^{-1}$, $\lambda = 0,6328 \text{ } \mu\text{m}$

$n_{\text{water}} = 1,33$ $v_{\text{water}} = 2,26 \times 10^8 \text{ m} \cdot \text{s}^{-1}$, $\lambda_{\text{water}} = 0,4758 \text{ } \mu\text{m}$

$n_{\text{glass}} = 1,5$ $v_{\text{glass}} = 2,00 \times 10^8 \text{ m} \cdot \text{s}^{-1}$, $\lambda_{\text{glass}} = 0,4219 \text{ } \mu\text{m}$



Fundamentals – complex refractive index

Definition and physical interpretation of refractive index

imaginary part - **absorption**:

Penetration depth:
$$d = \frac{1}{m_i k_0} = 0.16 \frac{\lambda}{m_i}$$

$$\begin{aligned} E &= E_0 e^{i(\omega t - m k_0 z + \phi)} \\ &= E_0 e^{i\omega t - i m_r k_0 z - m_i k_0 z + i\phi} \\ &= E_0 e^{-m_i k_0 z} e^{i(\omega t - m_r k_0 z + \phi)} \end{aligned}$$

Amplitude à z :

$$E_0(z) = E_0(z=0) e^{-m_i k_0 z}$$

Penetration depth d :

$$\begin{aligned} \frac{E_0(z=d)}{E_0(z=0)} &= e^{-1} \\ \Rightarrow d &= \frac{1}{m_i k_0} = 0.16 \frac{\lambda}{m_i} \end{aligned}$$

$$\lambda = 0.6328 \mu\text{m}$$

$$m_i = 0.1, d = 1 \mu\text{m}$$

$$m_i = 0.0001, d = 1 \text{ mm}$$

Fundamentals – balance of energy

Energy density and light intensity

Energy density (J/m^3): $u = \frac{1}{2} (\vec{E} \cdot \vec{D} + \vec{B} \cdot \vec{H})$

Poynting vector (W/m^2):

$$\vec{S} = \vec{E} \wedge \vec{H} = \frac{1}{2} \text{Re}(\vec{E} \wedge \vec{H}^*)$$

Complex function



In a homogeneous and isotropic medium: $\vec{H} = \frac{1}{\mu\omega} \vec{k} \wedge \vec{E}$

$$\vec{S} = \frac{1}{2} \text{Re}(\vec{E} \wedge \vec{H}^*) = \frac{1}{2\mu\omega} \text{Re}[\vec{E} \wedge (\vec{k} \wedge \vec{E}^*)] = \frac{1}{2\mu\omega} \|\vec{E}\|^2 \vec{k}$$

In an isotropic medium: $\vec{S} = v u \vec{n}$

Intensity: $I = \|\vec{S}\| \propto E^2$

Fundamentals – Polarization

Relation between incident wave and scattered wave
In far field

$$\begin{pmatrix} E_{\parallel s} \\ E_{\perp s} \end{pmatrix} = \frac{e^{ikr}}{-ikr} \begin{pmatrix} S_2 & S_3 \\ S_4 & S_1 \end{pmatrix} \begin{pmatrix} E_{\parallel i} \\ E_{\perp i} \end{pmatrix}$$

$S_3 = S_4 = 0$ for homogeneous or multilayered sphere

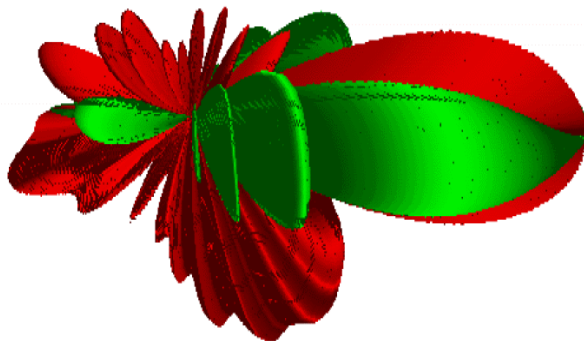
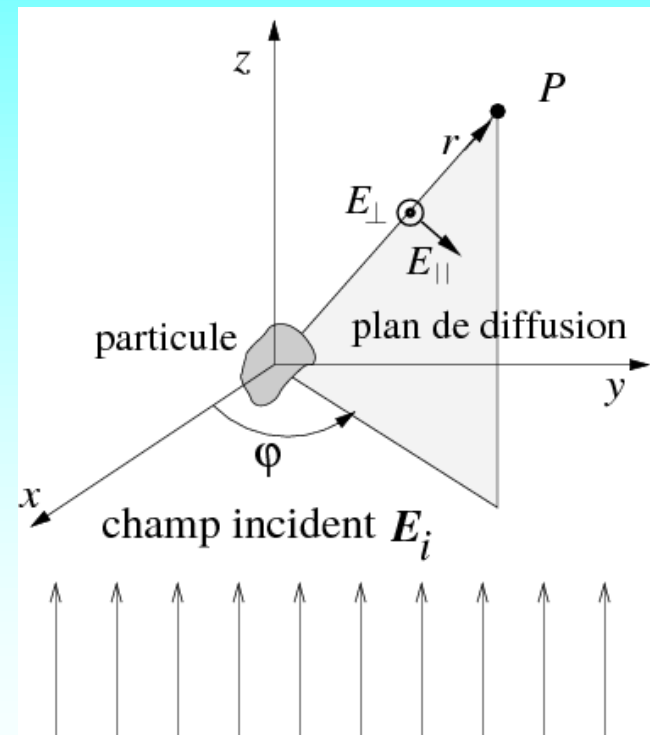


Diagram of a sphere illuminated by a Gaussian beam (off axis), perpendicular polarization in red and parallel in green



important

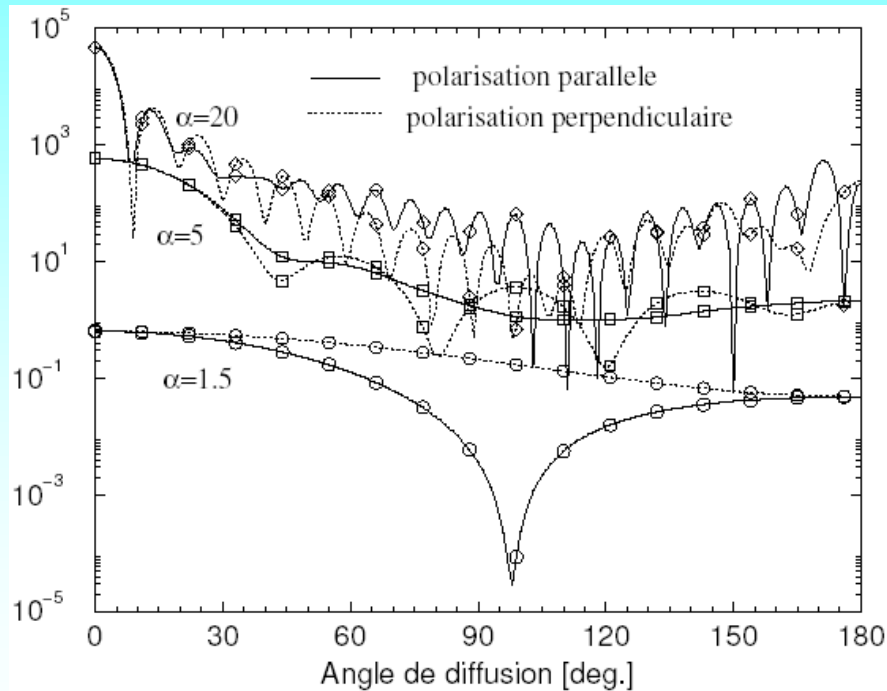
An exercise →

Fundamentals – scattering diagram

Angular distribution of intensity of the scattered wave
in far field

A sphere of diameter $d = \alpha\lambda/\pi$ and refractive index $m = 1.33$ illuminated by a plane wave

important



$$I(r, \theta, \varphi) = \frac{I_0 F(\theta, \varphi)}{k^2 r^2}$$

To be done:

We know S_j ($j=1, \dots, 4$) are only function of (θ, φ) in far field and $E_{\perp i} = E_0 p_{\perp} e^{i\omega t}$, $E_{\parallel i} = E_0 p_{\parallel} e^{i\omega t}$. Show this relation with scattering matrix according to $I(r, \theta, \varphi) = \|S\|$.

$$I_{\parallel}(\theta) = F(\theta, \varphi = 0) = |S_2|^2$$

$$I_{\perp}(\theta) = F(\theta, \varphi = 90^\circ) = |S_1|^2$$

Size parameter : $\alpha = \frac{\pi d}{\lambda}$

Fundamentals – Extinction, scattering, absorption

Integral properties of a scattering particle

Outside of the particle:

$$\mathbf{E} = \mathbf{E}_i + \mathbf{E}_s, \quad \mathbf{H} = \mathbf{H}_i + \mathbf{H}_s$$

Poynting vector :

$$\mathbf{S} = \frac{1}{2} \text{Re}\{\mathbf{E} \times \mathbf{H}^*\} = \mathbf{S}_i + \mathbf{S}_s + \mathbf{S}_{ext}$$

$$\mathbf{S}_i = \frac{1}{2} \text{Re}\{\mathbf{E}_i \times \mathbf{H}_i^*\}$$

$$\mathbf{S}_s = \frac{1}{2} \text{Re}\{\mathbf{E}_s \times \mathbf{H}_s^*\}$$

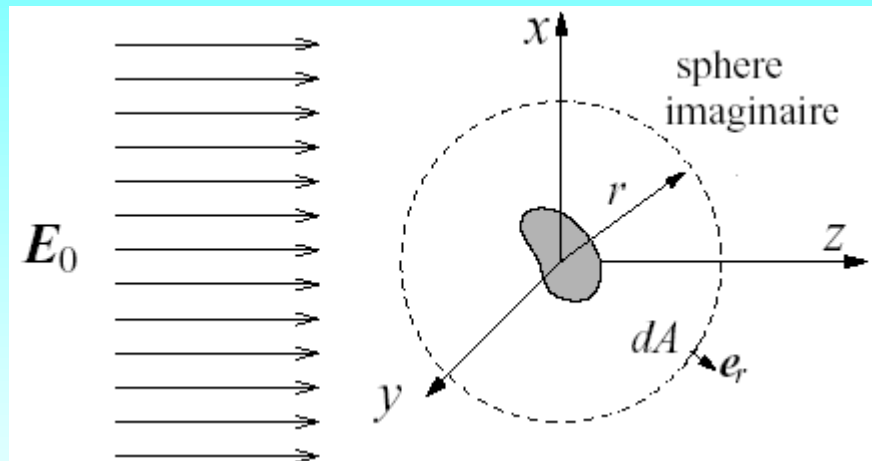
$$\mathbf{S}_{ext} = \frac{1}{2} \text{Re}\{\mathbf{E}_i \times \mathbf{H}_s^* + \mathbf{E}_s \times \mathbf{H}_i^*\}$$

Energy balance

$$W_a = - \int_A \mathbf{S} \cdot \mathbf{e}_r dA = W_i - W_s + W_{ext}$$

$$W_i = - \int_A \mathbf{S}_i \cdot \mathbf{e}_r dA, \quad W_s = \int_A \mathbf{S}_s \cdot \mathbf{e}_r dA, \quad W_{ext} = - \int_A \mathbf{S}_{ext} \cdot \mathbf{e}_r dA$$

$$W_{ext} = W_{abs} + W_{sca}$$



Fundamentals – Extinction, scattering, absorption

Definition of sections and factors

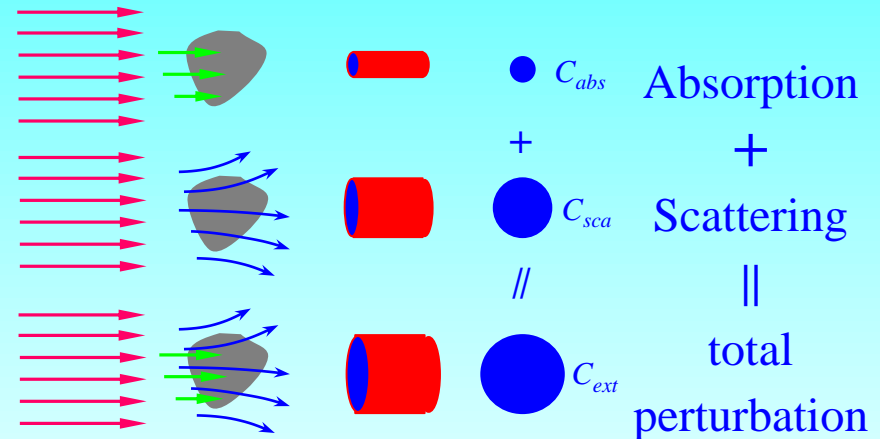
Efficiency sections:

Absorption sections: $C_{abs} = \frac{W_{abs}}{I_i}$

Scattering section: $C_{sca} = \frac{W_{sca}}{I_i}$

Extinction section : $C_{ext} = \frac{W_{ext}}{I_i}$

Physical interpretation



A is the geometrical section of the particle
area projected in the plan \perp incident direction.

Efficiency factors

$$Q_{ext} = \frac{C_{ext}}{A}, \quad Q_{abs} = \frac{C_{abs}}{A}, \quad Q_{sca} = \frac{C_{sca}}{A}$$

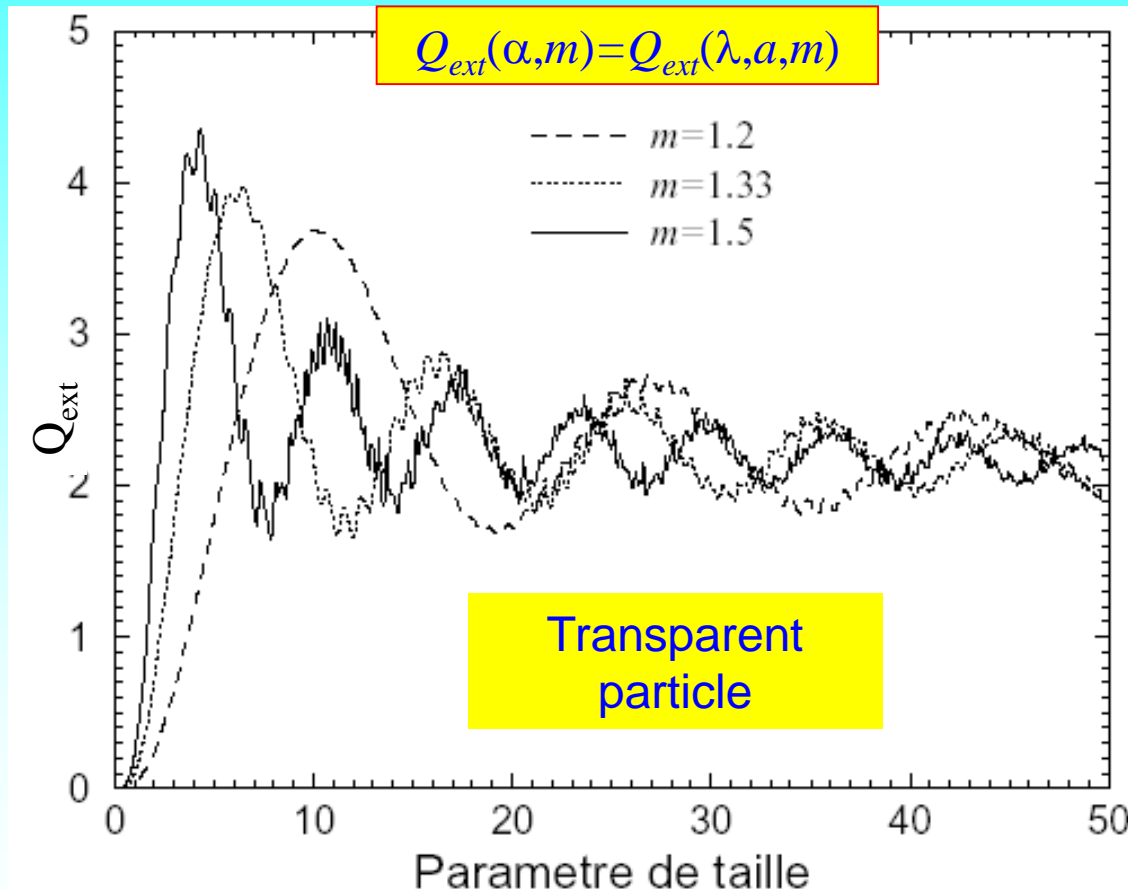
Transparent /no absorbing

$$C_{abs} = 0, \quad C_{ext} = C_{sca}$$

$$Q_{abs} = 0, \quad Q_{ext} = Q_{sca}$$

Fundamentals – Extinction, scattering, absorption

Read and understand the graphics



Small particle

$$Q_{ext} \sim \frac{d^4}{\lambda^4}$$

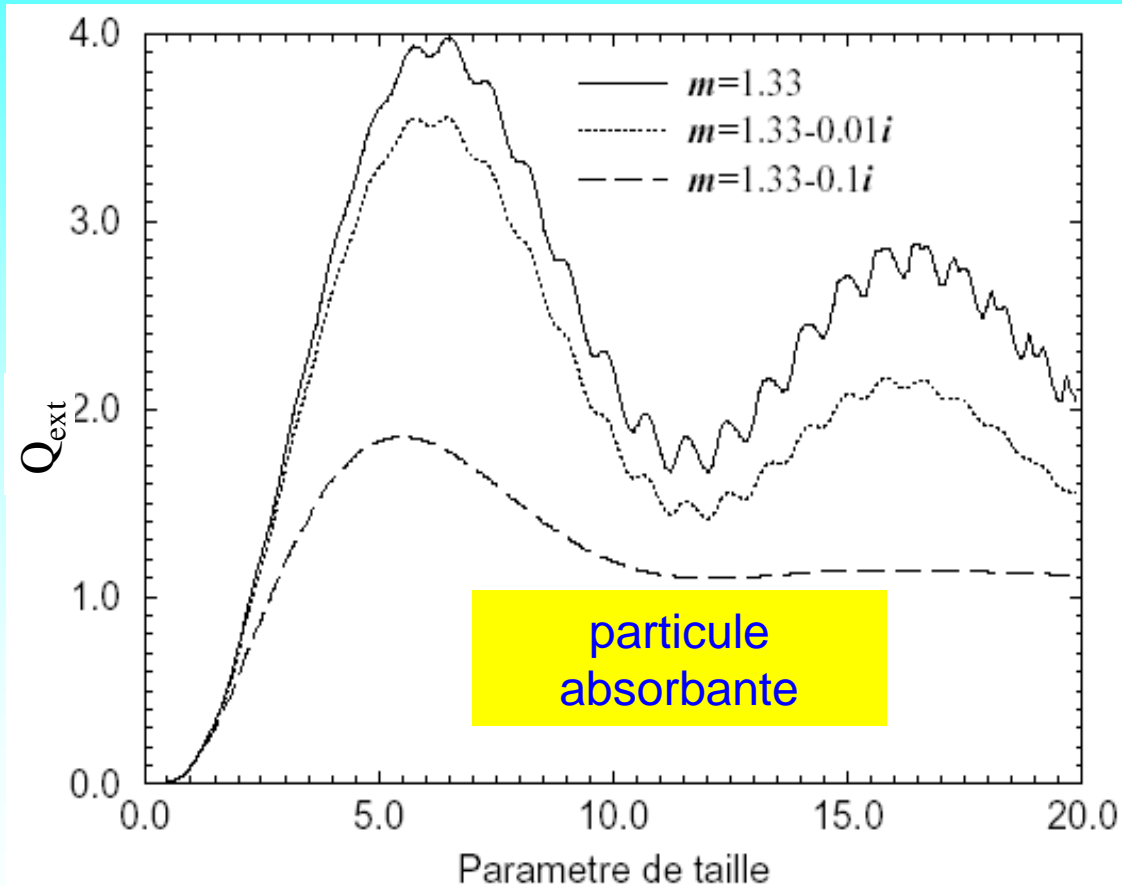
Big particle

$$Q_{ext} \rightarrow 2$$

Why is the sky blue and the sun red at the rising and at the set ?

Fundamentals – Extinction, scattering, absorption

Read and understand the graphics

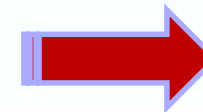


- The greater the absorption, the more smooth is the curve.
- The higher the absorption, the smaller the extinction factor.
- For small particles:

$$Q_{\text{ext}} \propto \alpha$$
- For big particles:

$$Q_{\text{ext}} \rightarrow 1$$

 i.e.: $C_{\text{ext}} \rightarrow A$

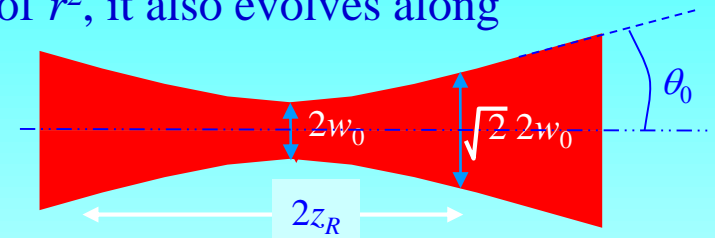


Fundamentals – Gaussian beam

Intensity distribution and geometric shape

- (a). **Intensity** decreases exponentially as function of r^2 , it also evolves along the beam axis.

$$I(r, z) = I_0 \left[\frac{w_0}{w(z)} \right]^2 \exp \left[-\frac{2r^2}{w^2(z)} \right]$$



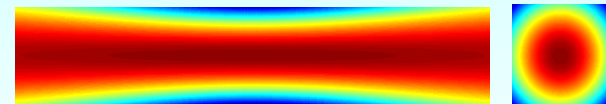
- (b). **Beam waist radius**

$$I(r = w(z)) = \frac{I(r = 0, z)}{e^2}$$

$$w(z) = w_0 \sqrt{1 + \left(\frac{\lambda z}{\pi w_0^2} \right)^2}$$

- (c). **Divergence of the beam**

$$\theta_0 = \arctan \left(\frac{\lambda}{\pi w_0} \right)$$



- (d). **Rayleigh length:** $z_R = \pi w_0^2 / \lambda$.

$$I(0, z_R) = \frac{I_0}{2}, \quad w(z_R) = \sqrt{2} w_0$$

Fundamentals – Gaussian beam

Some examples

Divergence:

1. $\lambda = 600 \text{ nm}$:

$w_0 = 10 \text{ }\mu\text{m}$: $z_R = 500 \text{ }\mu\text{m}$, $\theta = 0.02 \text{ rad}$

$w_0 = 1 \text{ mm}$: $z_R = 5 \text{ m}$, $\theta = 0.0002 \text{ rad}$

$w_0 = 1 \text{ cm}$: $z_R = 500 \text{ m}$, $\theta = 0.00002 \text{ rad}$

2. $w_0 = 100 \text{ }\mu\text{m}$:

$\lambda = 10.6 \text{ }\mu\text{m}$ (CO₂): $z_R = 3 \text{ mm}$, $\theta = 0.034 \text{ rad}$

$\lambda = 0.6328 \text{ }\mu\text{m}$ (He-Ne): $z_R = 5 \text{ cm}$, $\theta = 0.002 \text{ rad}$

$\lambda = 0.488 \text{ }\mu\text{m}$ (bleu YAG): $z_R = 6.4 \text{ cm}$, $\theta = 1.5 \text{ mrad}$

Intensity:

At the center:

$$I = I_0$$

On the border of the beam:

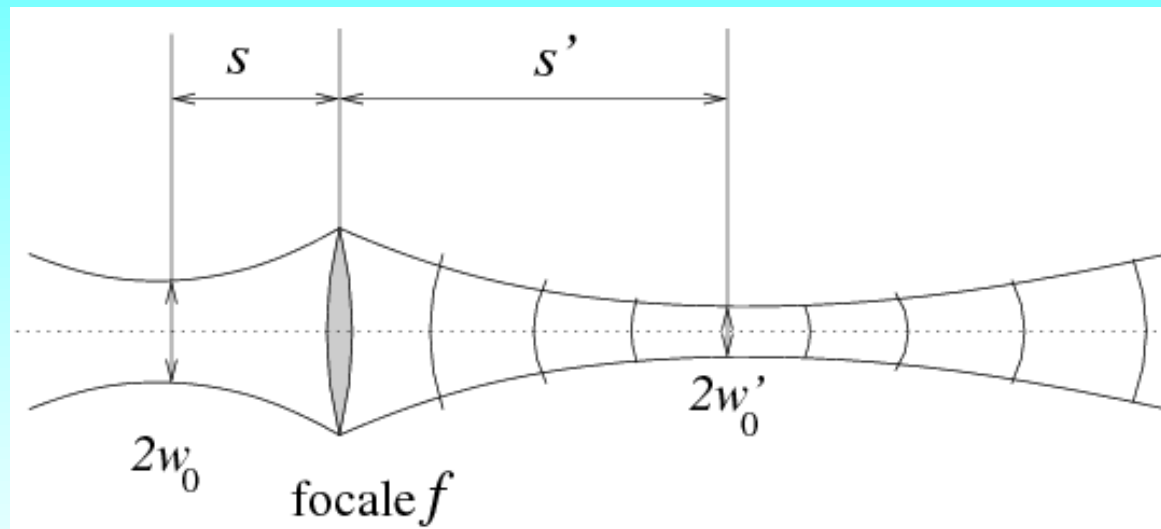
$$I(r = w_0) = \frac{I(r = 0, z = 0)}{e^2} = \frac{I_0}{e^2}$$

Total power of the beam:

$$\begin{aligned} P &= \int I(r, z) dS = \int_0^{2\pi} \int_0^\infty I(r, z) r dr d\theta \\ &= 2\pi I_0 \left[\frac{w_0}{w(z)} \right]^2 \int_0^\infty \exp\left[-\frac{2r^2}{w^2(z)} \right] r dr \\ &= I_0 \frac{\pi w_0^2}{2} \end{aligned}$$

Fundamentals – Gaussian beam

Thin lens equation
Position of the waists before and after a lens



$$\frac{1}{s'} - \frac{1}{s + \frac{z_R^2}{s + f'}} = \frac{1}{f'}$$

or

$$s' = f' \left[1 - \frac{f'(s + f')}{(s + f')^2 + z_R^2} \right]$$

Fundamentals – Gaussian beam

Thin lens equation

Position of the waists before and after a lens

Consequences :

- For quasi-parallel beams, z_R is much greater than all other distances, therefore $s'=f'$, behaviour of the geometrical optics.
- If the beam waist is in the object **focal plane** of a lens, the beam waist is at the image **focal plane** of the lens, since $s = -f \rightarrow s'=f'$. This result is radically different from the geometrical optics.
- The conjugate relation between s and s' depends not only on f' but also on z_R , hence on λ and w_0 , which is different from the geometrical optics

Fundamentals – Gaussian beam

Thin lens equation Position of the waists before and after a lens

- Beam waist ratio

$$m = \frac{w_0'}{w_0} = \left[\left(1 + \frac{s}{f'} \right)^2 + \left(\frac{z_R}{f'} \right)^2 \right]^{-1/2}$$

$$z_R' = m^2 z_R$$

1. In the case s tends to infinity, $m = 0 \rightarrow w' = 0$ (but **NO!** theoretical the limit of $w \sim \lambda/2$)
2. In the case $s = -f'$ (the waist is at the objet focal plan), so $m = f'/z_R = f'\lambda/\pi w_0^2$

The situation is very different from the GO!

$$\theta' = \lambda/\pi w_0' = w_0/f'$$

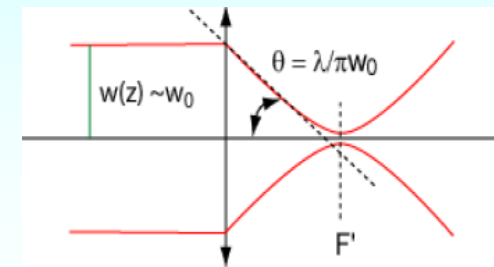
Example: $w_0 = 1 \text{ cm}$, $f' = 0.1 \text{ m}$, $\lambda = 0.6328 \mu\text{m} \rightarrow z_R = 500 \text{ m}$, $\theta = 0.001^\circ$

$s = -f' = -0.1 \text{ m} \rightarrow s' = 0.1 \text{ m}$, $w_0' = 2 \mu\text{m}$, $\theta' = 5.7^\circ$

We pass from a large waist to a small one ...

In practice, w_0 can be very small, so that w_0' is very large.

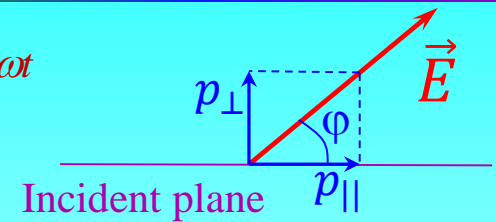
We find of course all the intermediate cases ...



Demonstration $I(r, \theta, \varphi)$:

$$E_{\perp i} = E_0 p_{\perp} e^{i\omega t}, \quad E_{\parallel i} = E_0 p_{\parallel} e^{i\omega t}$$

$$p_{\parallel}^2 + p_{\perp}^2 = 1$$



To simplify the problem, we show for a spherical particle:

$$\begin{pmatrix} E_{\parallel s} \\ E_{\perp s} \end{pmatrix} = \frac{e^{ikr}}{-ikr} \begin{pmatrix} S_2 & 0 \\ 0 & S_1 \end{pmatrix} \begin{pmatrix} E_{\parallel i} \\ E_{\perp i} \end{pmatrix}$$

$$\vec{E}_s = \frac{e^{ikr-i\omega t}}{-ikr} E_0 (S_2 p_{\parallel} \vec{e}_{\parallel} + S_1 p_{\perp} \vec{e}_{\perp})$$

$$I = \|\vec{S}\| = \frac{1}{2\mu\omega} \|\vec{E}\|^2 k = \frac{1}{2} \sqrt{\frac{\varepsilon}{\mu}} \frac{E_0^2}{(kr)^2} \left[(S_2 p_{\parallel})^2 + (S_1 p_{\perp})^2 \right] = \frac{I_0 F(\theta, \varphi)}{(kr)^2}$$

with

$$I_0 = \frac{E_0^2}{2} \sqrt{\frac{\varepsilon}{\mu}}$$

$$F(\theta, \varphi) = (S_2 p_{\parallel})^2 + (S_1 p_{\perp})^2$$

Example of question in exam:

A plane wave of amplitude E_0 polarized in the Oxz plane propagates along the axis Oz . It illuminates a homogeneous spherical particle. Express the scattered field E_{\perp}^s and E_{\parallel}^s in the far field as function of the scattering matrix elements S_1, S_2, S_3, S_4 and the observation distance r for the following cases:

- the observation point is in the Oxz plane,
- the observation point is in the Oyz plane,
- the observation point is in the plane containing the Oz axis and makes an angle $\psi = 30^\circ$ relative to the Ox axis.

(a) Parallel polarization:

$$E_{\parallel}^s = \frac{e^{ikr}}{-ikr} E_0 S_2$$

(b) Perpendicular polarization:

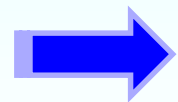
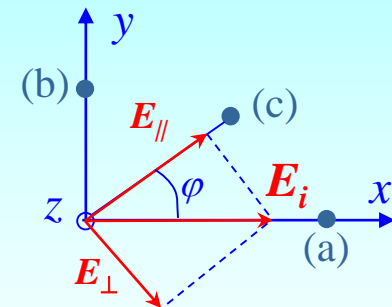
$$E_{\perp}^s = \frac{e^{ikr}}{-ikr} E_0 S_1$$

(c) 2 components:

$$p_{\parallel} = \cos 30^\circ = \frac{\sqrt{3}}{2}, \quad p_{\perp} = \sin 30^\circ = \frac{1}{2}$$

$$E_{\parallel}^s = \frac{e^{ikr}}{-ikr} E_{\parallel}^s S_2 = \frac{e^{ikr}}{-ikr} \frac{\sqrt{3}}{2} E_0 S_2$$

$$E_{\perp}^s = \frac{e^{ikr}}{-ikr} E_{\perp}^s S_1 = \frac{e^{ikr}}{-ikr} \frac{1}{2} E_0 S_1$$



Exercise on extinction

A cloud of water drops (refractive index $m=1.333$) of thickness $L=10$ mm is illuminated by a light of two wavelengths $\lambda=0.6328$ μm and 0.450 μm .

1. We assume that the drops are monodisperse of diameter $d=100$ nm. What are the extinction coefficients of particles at the two wavelengths?
2. The measured transmittance I/I_0 at $\lambda=0.6328$ is 0.8. Deduce the concentration of particles. What is then the transmittance at $\lambda=0.450$ μm ?
3. Next, we suppose that the environment is composed of two populations of drops with diameters of 100 nm and 1000 nm with a concentration of 10^{17} m^{-3} and 10^{13} m^{-3} respectively. What is the transmittance of the medium?

$$\alpha_1=0.5, \alpha_2=0.7 (\alpha_2/\alpha_1)^4=3.8$$

1. $d \ll \lambda$, so Rayleigh approximation: $Q_{ext} = \frac{8\alpha^4}{3} \Re \left(\frac{m^2 - 1}{m^2 + 2} \right) = 0,0333, 0.128$
2. Beer law: $N = \frac{\ln(I_0/I)}{LC_{ext}} = \frac{\ln(1/0,8)}{10000 \times 0,0333 \times \pi 0,05^2} = 0.0085 \mu\text{m}^{-3} = 8,5 \times 10^{15} \text{m}^{-3}$
 $I/I_0=0.8^{3.8}=0.42$
3. $Q_{ext}(100\text{nm})=0.0333, Q_{ext}(1\mu\text{m})=3.6$ (see figure)

$$\frac{I}{I_0} = \exp[-L(N_1 C_{ext,1} + N_2 C_{ext,2})] = 0,58 = 58\%$$

$$\begin{aligned} & \exp[-0.01 \times 10^{17} \times 10^{-16} \pi (50^2 \times 0.033 + 5^2 \times 3.6)] \\ & = \exp[-0.001 \times 3.14 (83 + 90)] = \exp(-0.54) \end{aligned}$$